

CLUES to overcome Biases

Jenny Sorce

CLUES workshop

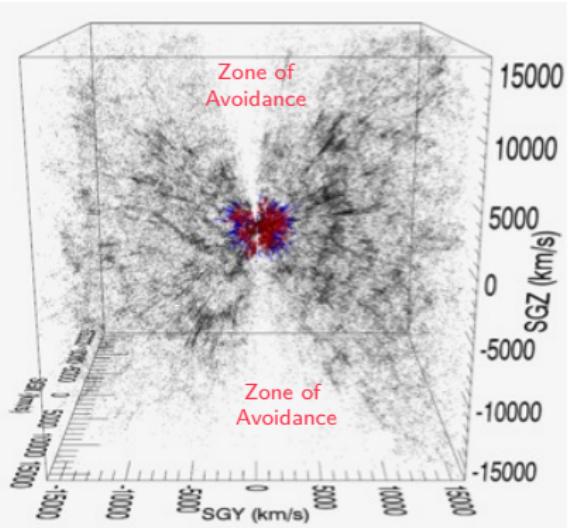
Potsdam - August, 2014

IPNL / University of Lyon - AIP / University of Potsdam

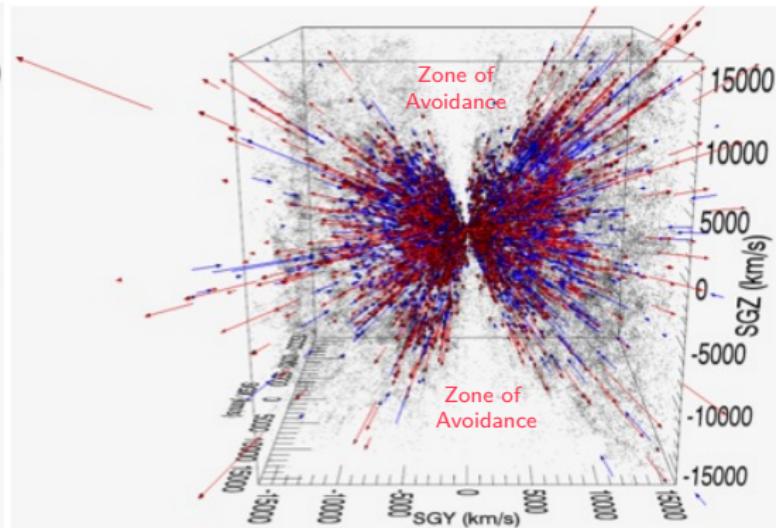
Collaborators: Helene Courtois, Stefan Gottlöber, Yehuda Hoffman, Daniel Pomarède, Matthias Steinmetz, Brent Tully, Gustavo Yepes

CF1 and CF2: Evolution of Cosmicflows Catalogs

Galaxies with measured radial peculiar velocities



Cosmicflows-1
about 2000 constraints
Tully et al. 2008



Cosmicflows-2
about 8000 constraints
Tully et al. 2013

With each catalog, improvements of :

- **quality** (e.g. accuracy)
- **quantity** (e.g. farther, *Zone of Avoidance*)

Black dots: XSCZ redshift catalog

How Catalogs Are Built

$$v_{obs} = H_0 \times d + v_{pec\ radial} \quad (1)$$

↓ redshift ↓ constant ↓ distance ↓ velocity

$$m - M = 5 \log_{10}(d(Mpc)) + 25 \quad (2)$$

$$\begin{cases} m \leftrightarrow \text{Photometry} \\ M \leftrightarrow \text{ITFR} : L \propto v_{HI}^{\alpha} \end{cases}$$

Mainly (Tully & Fisher 1977)

Observations

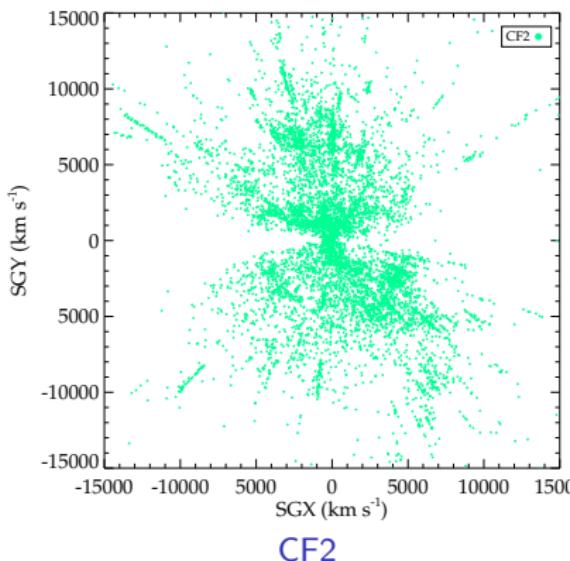
Calibrations

$$(2) \rightarrow d \xrightarrow{(1)} \text{radial } v_{pec} \rightarrow \text{Cosmicflows Catalogs} \rightarrow \text{e.g. CF1, CF2}$$



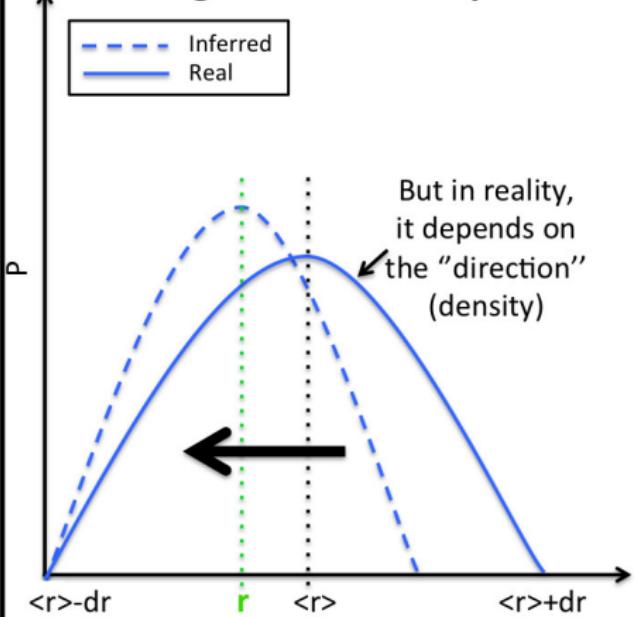
Constrained Simulations

methods, mocks, observational data

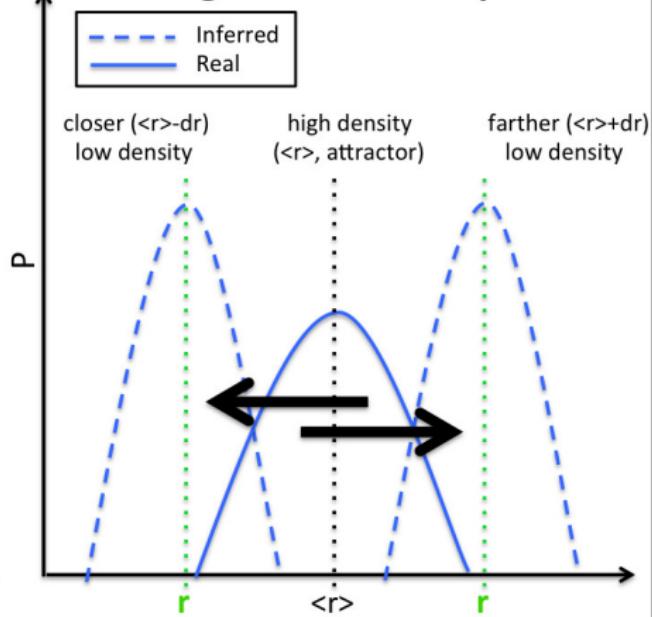


Deeper Catalogs CF2: Stronger Biases' Effects

Homogeneous Malmquist Bias

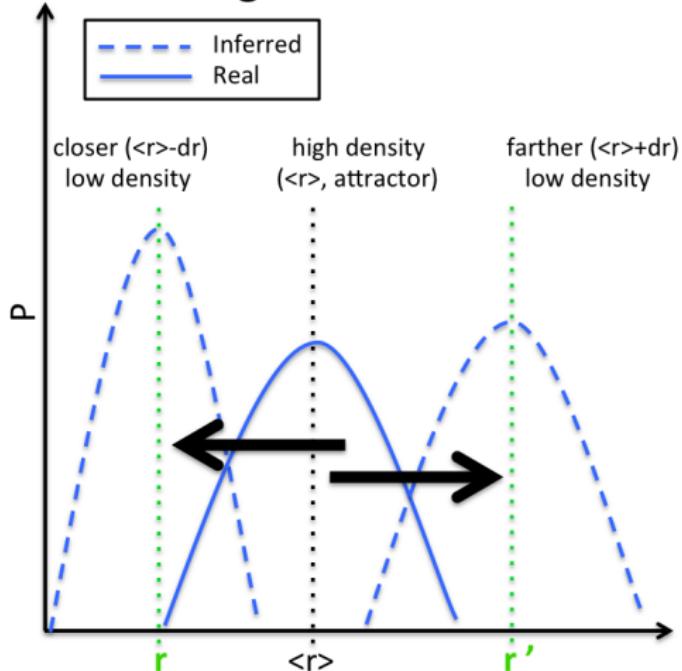


Inhomogeneous Malmquist Bias



Deeper Catalogs CF2: Stronger Biases' Effects

Error Lognormal Distribution



Error Lognormal Distribution:

Example:

$$\mu = 35 \pm 0.8 \text{ mag} \quad (20\% \text{ on } d) \Rightarrow$$

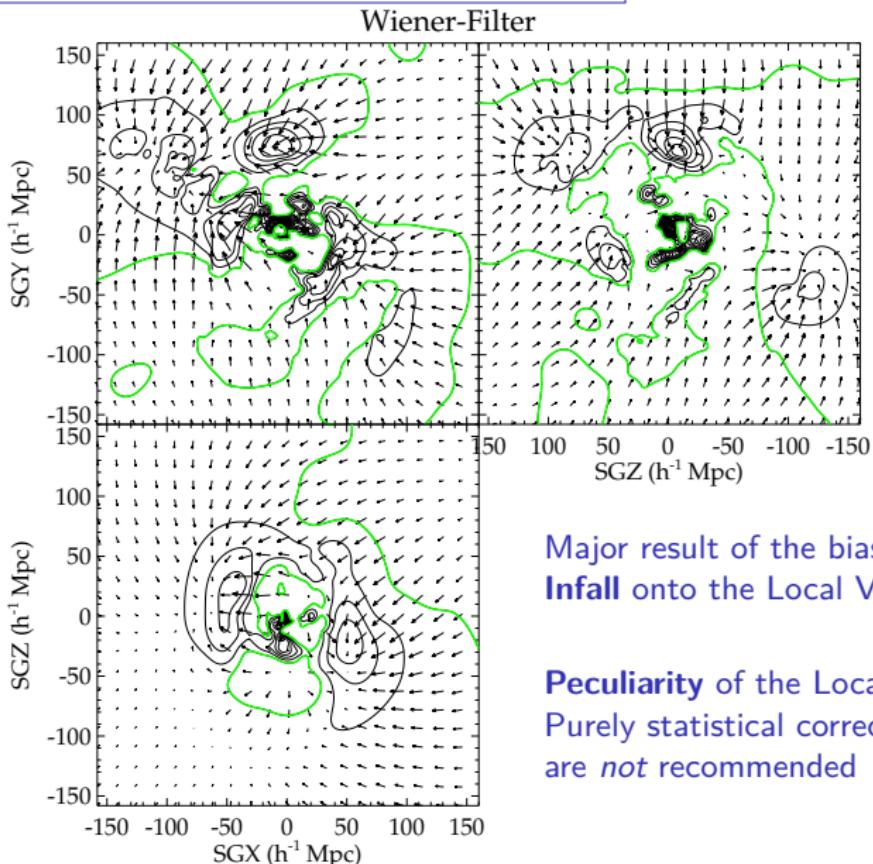
$$d = 100 [69; 145] \text{ Mpc} \quad (-31; +45)$$

$$H_0 = 75 \text{ km s}^{-1} \text{ Mpc}^{-1} \text{ & } v_{obs} = 7500 \text{ km s}^{-1} \Rightarrow$$

$$v_{pec} = 0 \quad [2311; -3338] \text{ km s}^{-1}$$

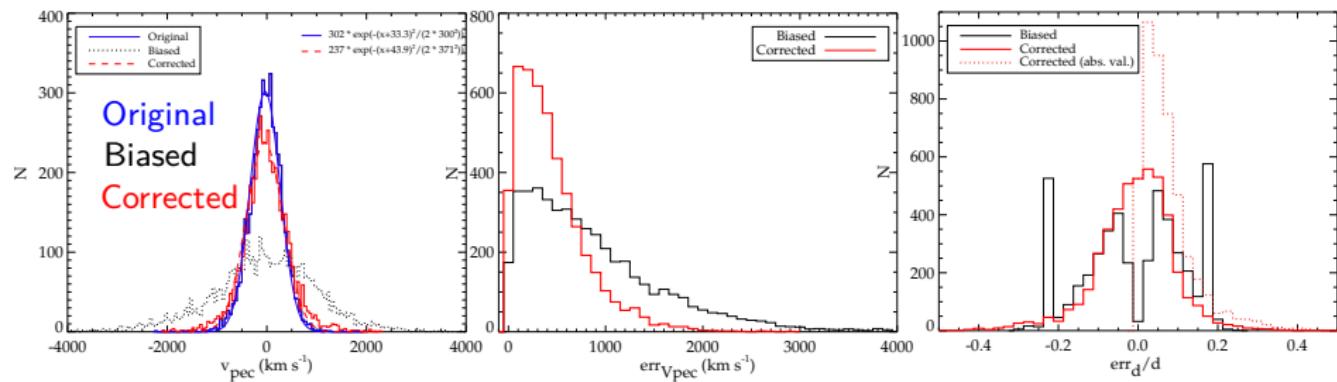
Asymmetry

General Infall onto the Local Volume



Proposition of a Method to Minimize the (Asymmetry) Bias(es)

Method 1: use of CF2-Mocks



$$\text{if } v_{\text{pec}} > 0, \quad v_{\text{pec c}} = (1 - w)[p(v_{\text{pec}} - \sigma_{v_{\text{pec}}}) + (1 - p)(v_{\text{pec}} + \sigma_{v_{\text{pec}}})] + w v_{\text{pec}}$$

$$\text{if } v_{\text{pec}} < 0, \quad v_{\text{pec c}} = (1 - w)[p(v_{\text{pec}} + \sigma_{v_{\text{pec}}}) + (1 - p)(v_{\text{pec}} - \sigma_{v_{\text{pec}}})] + w v_{\text{pec}}$$

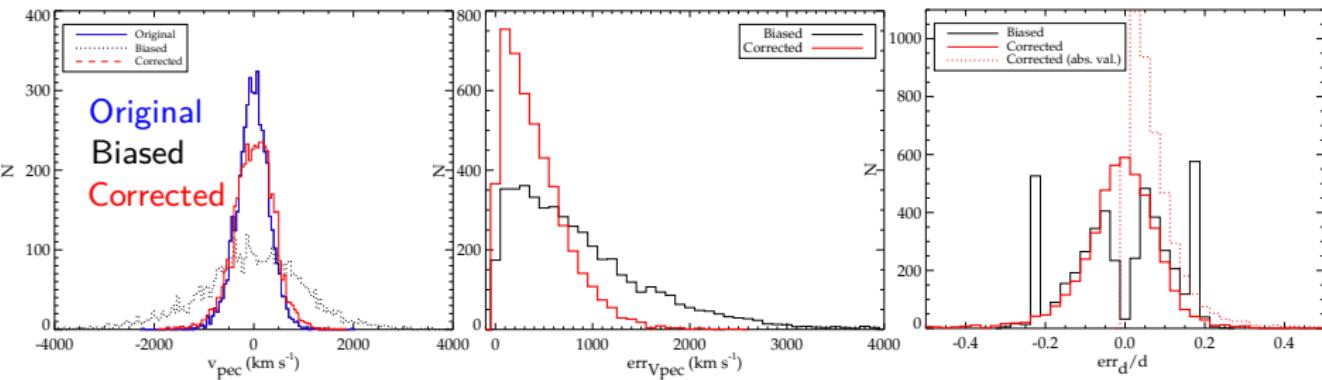
- p : probability $v_{\text{pec}} \notin$ theoretical Gaussian (from the mock) (Sheth and Diaferio, 2001)
- w : weighted uncertainty on v_{pec}

After correction:

- distances computed accordingly: $d_c = (v_{\text{obs}} - v_{\text{pec c}})/H_0$
- 5-6% fractional error on distances assumed.

Proposition of a Method to Minimize the Infall (Biases)

Method 2: use of the Wiener-Filter reconstructed velocity field



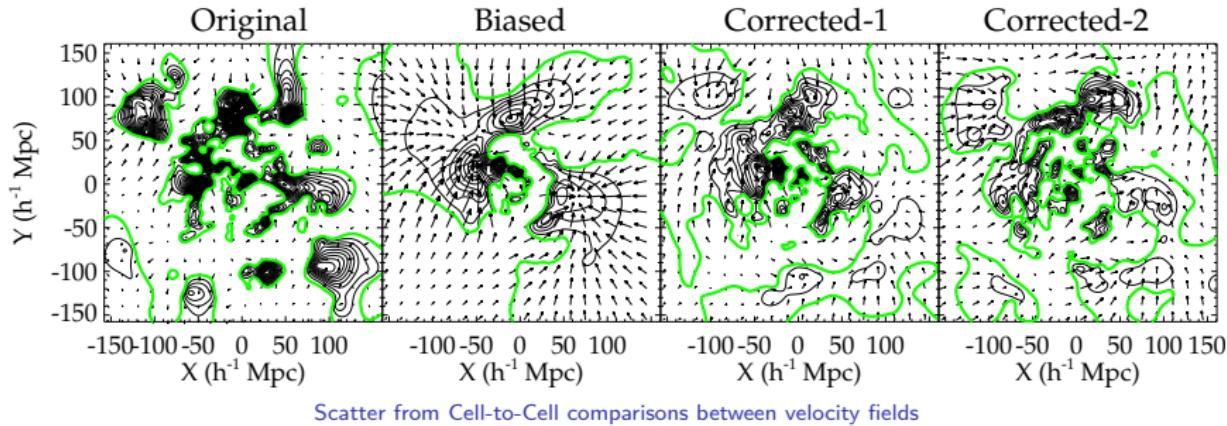
$$v_{\text{pec } c} = g \times v_{\text{pec rad}}^{\text{WF}} + (1 - g) \times v_{\text{pec}}$$

- g : chosen to minimize the monopole term ($\propto \vec{\nabla} \cdot \vec{v}$) at $150 \text{ h}^{-1} \text{ Mpc}$ of the Wiener-Filter velocity field obtained after correction

After correction:

- distances computed accordingly: $d_c = (v_{\text{obs}} - v_{\text{pec } c})/H_0$
- 5-6 % fractional error on distances assumed.

Confirming Bias Correction on Mocks



| Mock | σ_{WF} km s ⁻¹ | σ_{WF} h ⁻¹ Mpc | σ_{simu} km s ⁻¹ | σ_{simu} h ⁻¹ Mpc | $\sigma_{WF \; div}$ km s ⁻¹ | $\sigma_{WF \; div}$ h ⁻¹ Mpc |
|-------------|-------------------------------------|--------------------------------------|---------------------------------------|--|--|---|
| Biased | 182 | 3.8 | 220 | 4.5 | 195 | 4. |
| Corrected-1 | 116 | 2.4 | 149 | 3.1 | 129 | 2.7 |
| Corrected-2 | 156 | 3.2 | 171 | 3.5 | 156 | 3.2 |

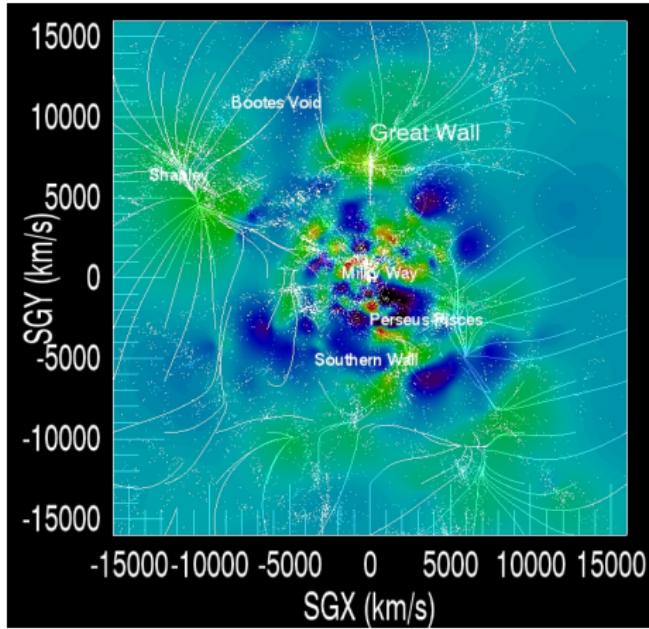
Comparison between methods:

- Full: $\sigma = 92 \text{ km s}^{-1}(1.9 \text{ h}^{-1} \text{ Mpc}) < 2 \text{ h}^{-1} \text{ Mpc}$
- Div.: $\sigma = 67 \text{ km s}^{-1}(1.4 \text{ h}^{-1} \text{ Mpc})$ (linear theory)

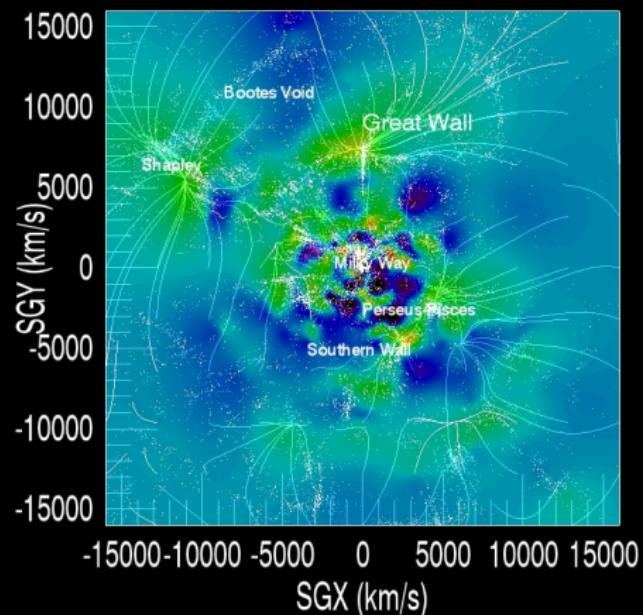
Large Scale Structures, voids and flows of the reference simulation are robustly recovered with both methods at levels close to 3-4 h⁻¹ Mpc .

Confirming Bias Correction with CF2

Method 1



Method 2

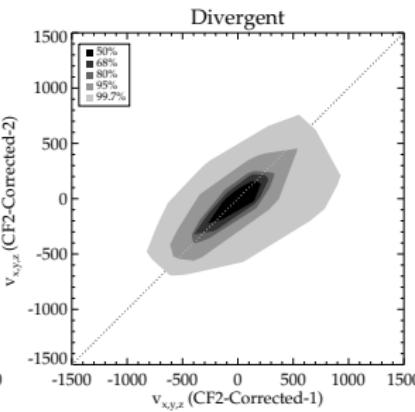
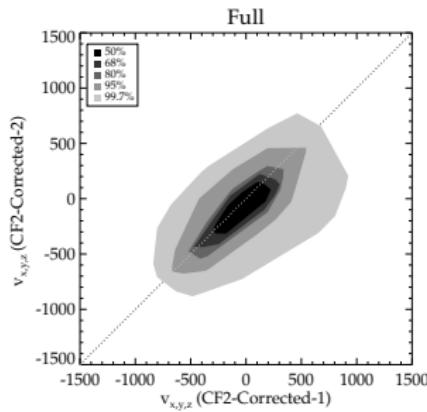


The Large Scale Structures and voids of the Local Universe are robustly recovered with both methods

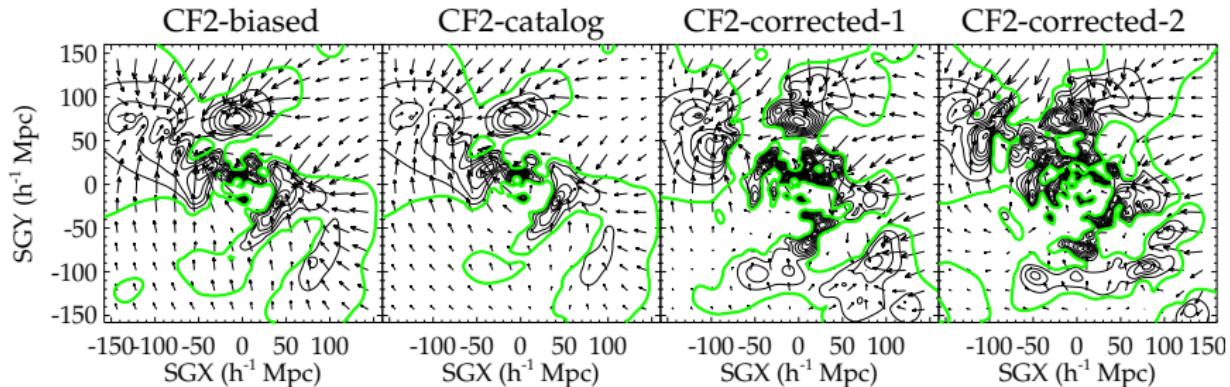
White Dots = XSCZ redshift catalog

SDvision Software (Daniel Pomarède)

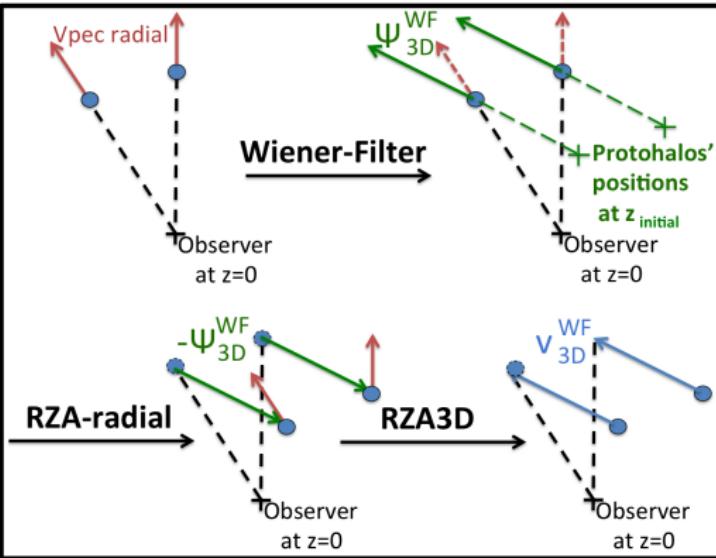
Consistency of the Two Corrective Methods



Scatter in terms of velocity: 58 / 48 km s⁻¹
 leading to a scatter in terms of displacement:
1.2 / 1 h⁻¹ Mpc
 $< 3 \text{ h}^{-1} \text{ Mpc}$ (linear theory + bias residual)



Reverse Zel'dovich Approximation



Reconstructions

Wiener-Filter (Λ CDM) (Zaroubi et al. 1995)



RZA3D (Doumler et al. 2013a,b,c ; Sorce et al. 2014)



Constrained Realizations (Λ CDM)

(Hoffman & Ribak 1991)



Initial Conditions



Constrained Simulations

Reverse Zel'dovich Approximation:

$$\vec{x}_{\text{init}}^{\text{RZA}} = \vec{r} - \frac{\vec{v}}{H_0 f(t_{\text{init}})}$$

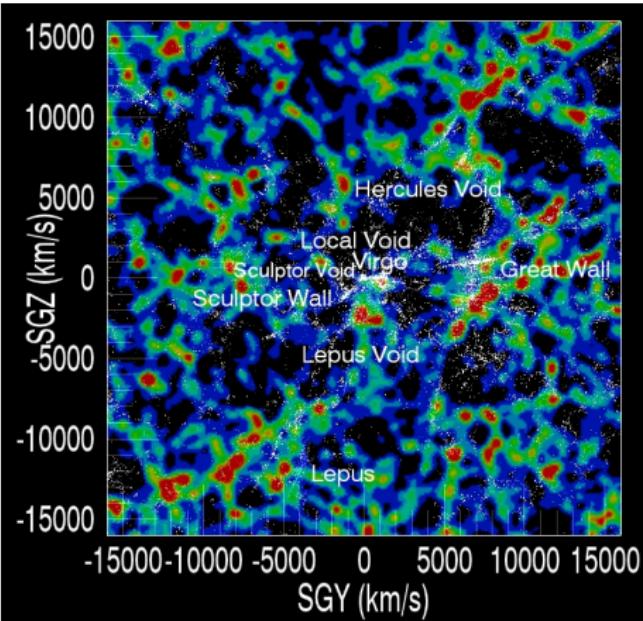
growth rate : $f(t) = \frac{d(\ln D(t))}{d(\ln a(t))}$ growth factor
scale factor

Linear Theory at 1st order valid down to 2 h^{-1} Mpc

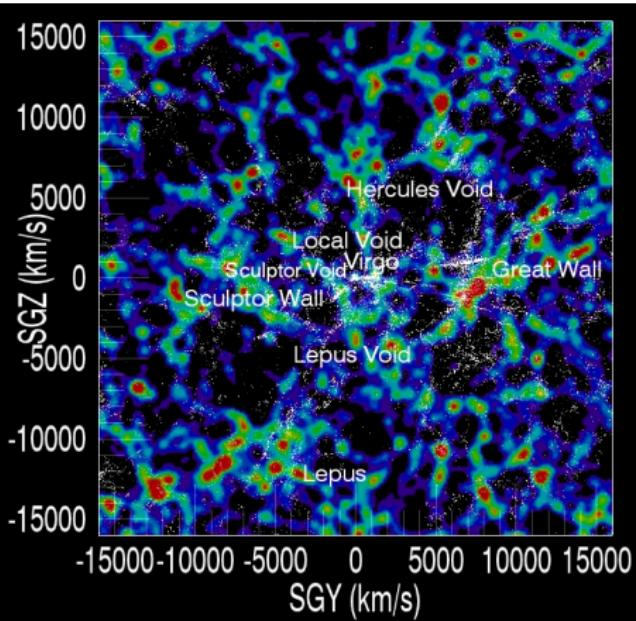
Formation of Structures: A Movie

At $z = 0$,

Method 1

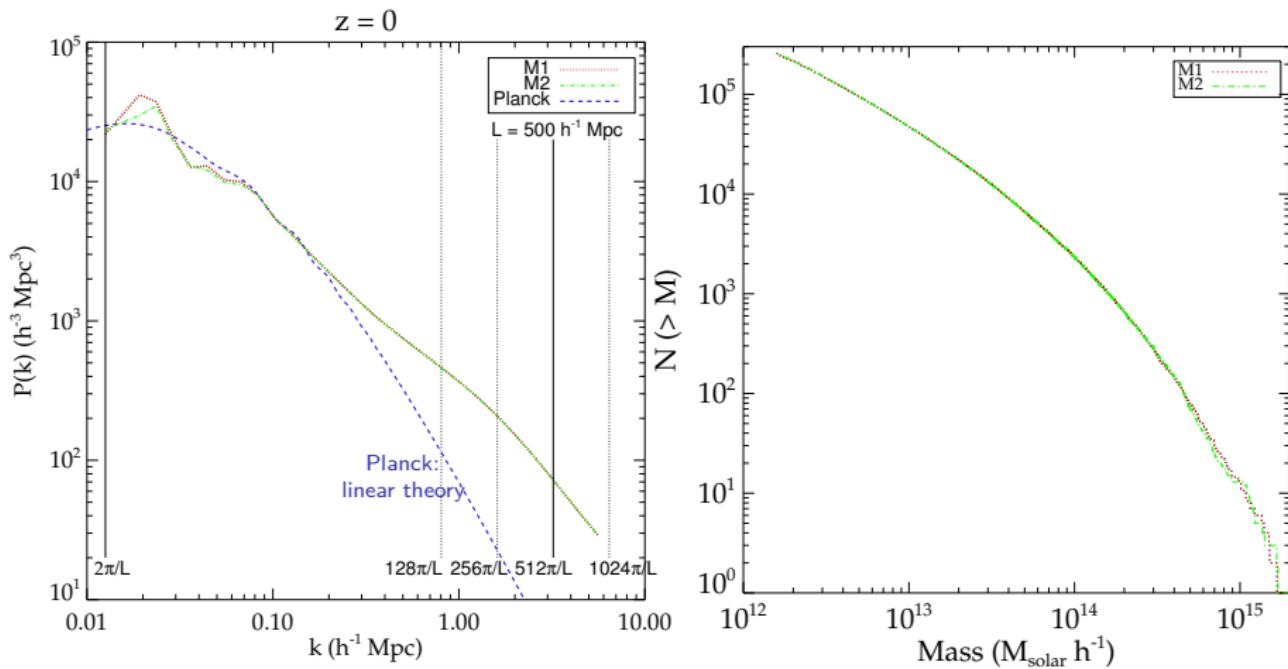


Method 2



Periodic Boundary Conditions, $L=500 \text{ } h^{-1} \text{ Mpc}$, $n^3=512^3$

Validation of Second Generation Constrained Simulations



Replicas of Virgo

Method 1 correction $\approx 0.5 \sigma$

Method 2 correction $\approx 2 \sigma \rightarrow$ less credit to observational measurement

| Virgo | SGX | SGY | SGZ | $v_{pec\ rad}$ | v_{pec} | d | M | δ_x | δ_y | δ_z | δ_{tot} |
|-------------------|-------|-------|--------|----------------|-----------|------|-----------------|------------|------------|------------|----------------|
| Observed | -2.80 | 12.0 | -0.561 | 207 | / | 12.3 | 40 [†] | | | | |
| Corrected M1 | -2.79 | 11.9 | -0.559 | 207 | | 12.3 | / | | | | |
| Inserted in CR M1 | 2.82 | 9.90 | -1.02 | / | 311 | | / | | | | |
| Corrected M2 | -2.73 | 11.7 | -0.547 | 232 | | 11.4 | / | | | | |
| Inserted in CR M2 | -1.46 | 9.63 | 4.70 | / | 280 | | / | | | | |
| Simulated M1 | -6.61 | 10.93 | -2.42 | 224 | 472 | 13.0 | 1.8 | -3.81 | -1.05 | -1.86 | 4.37 |
| | -4.15 | 9.02 | 2.64 | 145 | 562 | 10.3 | 1.4 | -1.34 | -2.95 | 3.21 | 4.56 |
| | -5.43 | 9.61 | -3.74 | 345 | 489 | 11.7 | 1.0 | -2.63 | -2.36 | -3.17 | 4.75 |
| | 0.27 | 11.16 | -1.15 | 48 | 440 | 11.2 | 4.9 | 3.07 | -0.82 | -0.59 | 3.23 |
| σ | 3.01 | 1.03 | 2.75 | 125 | 51 | 1.13 | 1.8 | 3.01 | 1.03 | 2.75 | 0.68 |
| Simulated M2 | -3.42 | 10.3 | -4.93 | 237 | 311 | 11.9 | 1.1 | -0.69 | -1.41 | -4.38 | 4.66 |
| | -3.18 | 7.18 | -1.28 | 392 | 421 | 7.96 | 1.4 | -0.45 | -4.52 | -0.73 | 4.60 |
| | -5.57 | 13.5 | 1.61 | 291 | 385 | 14.7 | 1.7 | -2.84 | 1.48 | 2.16 | 3.98 |
| | -5.99 | 14.0 | 0.334 | 277 | 384 | 15.2 | 1.7 | -3.26 | 1.99 | 0.88 | 4.07 |
| σ | 1.45 | 3.15 | 2.83 | 66 | 46 | 3.31 | 0.30 | 1.45 | 3.15 | 2.83 | 0.35 |

Robust Large-Scale Environment

Method 1
(five simulations)

Method 2
(five simulations)

Virgo region ?

Virgo region ?

Major Attractors and Voids of the Local Universe are **simulated repetitively**
↪ **robust Large-Scale Environment**
to study local objects

Robust Large-Scale Environment

Method 1
(five simulations)

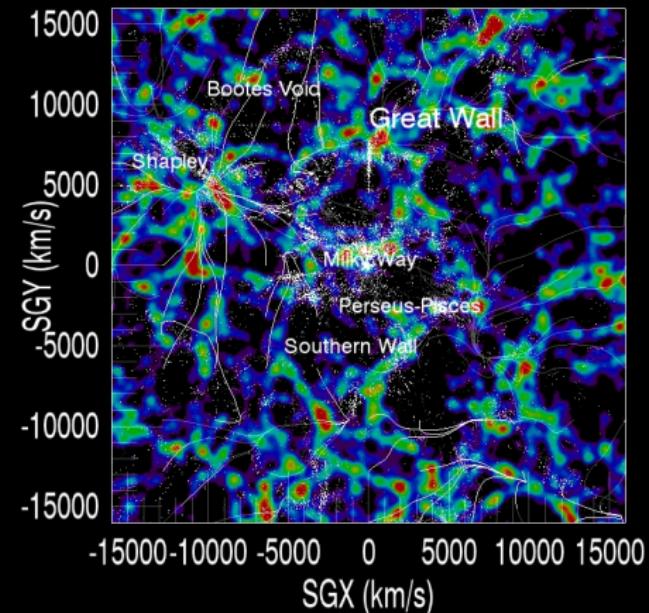
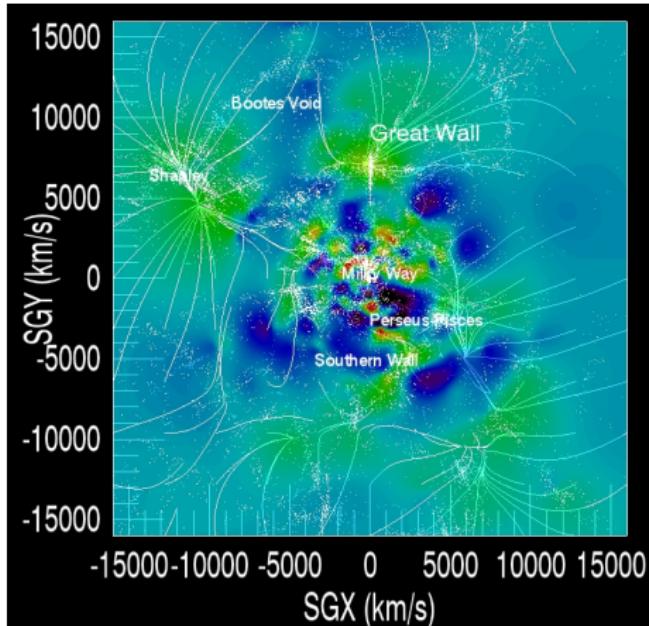
Virgo region ✓

Method 2
(five simulations)

Virgo region X

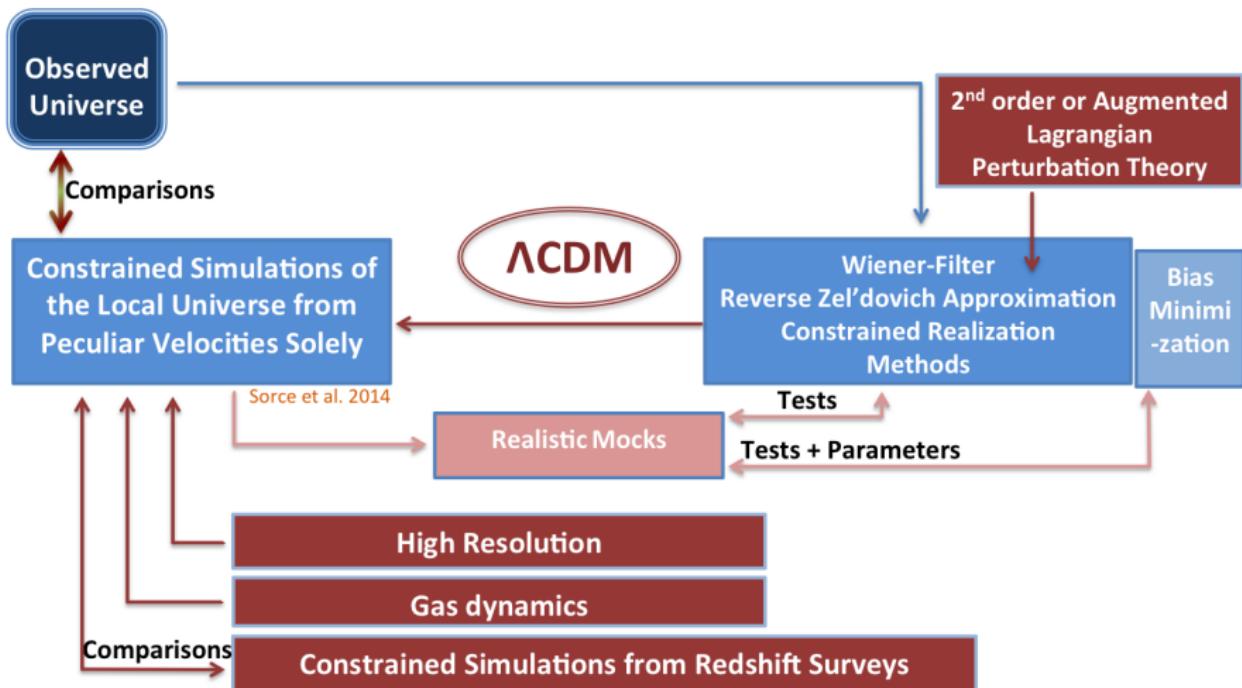
Major Attractors and Voids of the Local Universe are **simulated repetitively**
↪ **robust Large-Scale Environment**
to study local objects

Comparisons Observations/Reconstructions/Simulations



Observation/Reconstruction/Simulation

Overview and Prospectives



Acknowledgements

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